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研究課題名(和文) Can planetesimal accretion break planet resonance?

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研究代表者

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研究成果の概要(和文)：系外惑星にどのような傾向があるかを調べるために、ニューラルネットワークを使用した。元々は微惑星の降着によって惑星半径が増大するかどうかを調べるために設計したが、観測で直接測定することができない惑星の特性を推定することに、より威力を発揮した。2番目の論文では、ウェブブラウザで実行することのできるシンプルな気候解析プログラムを開発した。これによって、ユーザーは地球に似た系外惑星のほんのわずかな違いが全く違う環境につながることを体験することが出来る。3番目の論文ではニューラルネットワークによる画像識別を用いることで、火星表面の自動クレーター計数ソフトを開発した。

研究成果の学術的意義や社会的意義

While many exoplanets have been found, the measured properties of individual planets is small. This situation will not change, as observation techniques are sensitive to particular properties and orbits. To understand the exoplanet catalogue, techniques must be developed to impute missing values.

研究成果の概要(英文)：This research explored the use of neural networks to identify trends in the exoplanet archive. Originally designed to explore the increase in planetary radius due to accretion, greater success was achieved with imputing missing properties (including planet mass and radius) based on the multidimensional density function found by the neural network. Neural networks excel at finding trends in multidimensional data that are difficult to identify through plotting or other means. Such interpolation is necessary for properties that cannot be measured through observations. A second paper developed a simple climate code that could run on a web-browser to allow users to explore how small changes to an Earth-like planet could have a major impact on the environment. This also utilised a neural network to render the planet and bring it to a wider audience. A third paper used the image identification of neural networks to develop an automatic crater counter on the surface of Mars.

研究分野：Astrophysics

キーワード：catalogue exoplanet planet formation neural networks machine learning astrophysics planetary science

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様式 C - 19、F - 19 - 1、Z - 19 (共通)

1 . 研究開始当初の背景

Since the early 1990s, thousands of exoplanets have been discovered. Yet despite this high number, only a few properties about each planet are measured. This is a consequence of the detection techniques, which are sensitive to different planet properties and orbits. For example, a planet on a longer orbit will be less likely to transit the star, making a measurement of the planet's radius impossible. This situation will therefore not improve with more observations and the exoplanet catalogue will remain large but sparsely populated with measured properties. This has made utilising the exoplanet archive difficult, presenting challenges for comparing planet formation theories against the collected data.

2 . 研究の目的

The original goal of this project was to use numerical simulations to supplement missing observations and explore a potential pattern between planet radius and a history bombardment. This was less successful than planned, due to the difficulty with modelling the planet atmosphere sufficiently well to constrain a small increase in radius. We therefore moved to consider the use of neural networks to find patterns in the properties listed in the exoplanet archive. The goal was to see if missing planet properties could be estimated by the network and if these estimations could reveal information about the planet population that could be used to better understand planetary formation and evolution.

3 . 研究の方法

Neural networks excel at identifying patterns in high dimensional data. They have been traditionally used in image processing tasks and –more recently—in exoplanet science to identify the location of a transiting planet in noisy data. This was the first time that a neural network has been applied to the problem of generating missing planet properties.

The network is trained on planets with a set of properties that all have measured values. The network generates a probability distribution in multi-dimensional space that represents the known properties of the planets. By drawing from this space, the network can generate likely values for missing planet properties, based on the known measurements. Repeatedly drawing from the probability distribution creates a distribution for the missing property that can both estimate the value but also inform on that region of the parameter space. For example, a bimodal distribution for a missing mass value indicates two types of planets can be found with similar periods and radii.

4 . 研究成果

(1) The neural network was able to estimate the mass of a planet observed with the radial velocity technique (measured minimum mass and no radius) with an average error of 1.5 the true value. The radius of these planets could be estimated with an average error of 1.4 the true value. For planets observed with the transit technique where the radius is measured but there is no mass, the planet mass could be estimated to within 2.7 of the true value. This research was published in the *Astrophysical Journal*.

(2) In addition to estimating a single value for the missing property, a distribution of possible values could be calculated and used to inform on trends within the exoplanet catalogue. Bimodal distributions in mass or radius indicate that two different populations of planets commonly exist around stars with similar masses on similar orbits. For example, the neural network predicted that the mass and radius of WASP-8b could be either a hot Jupiter or super Earth (see figure 1 below).

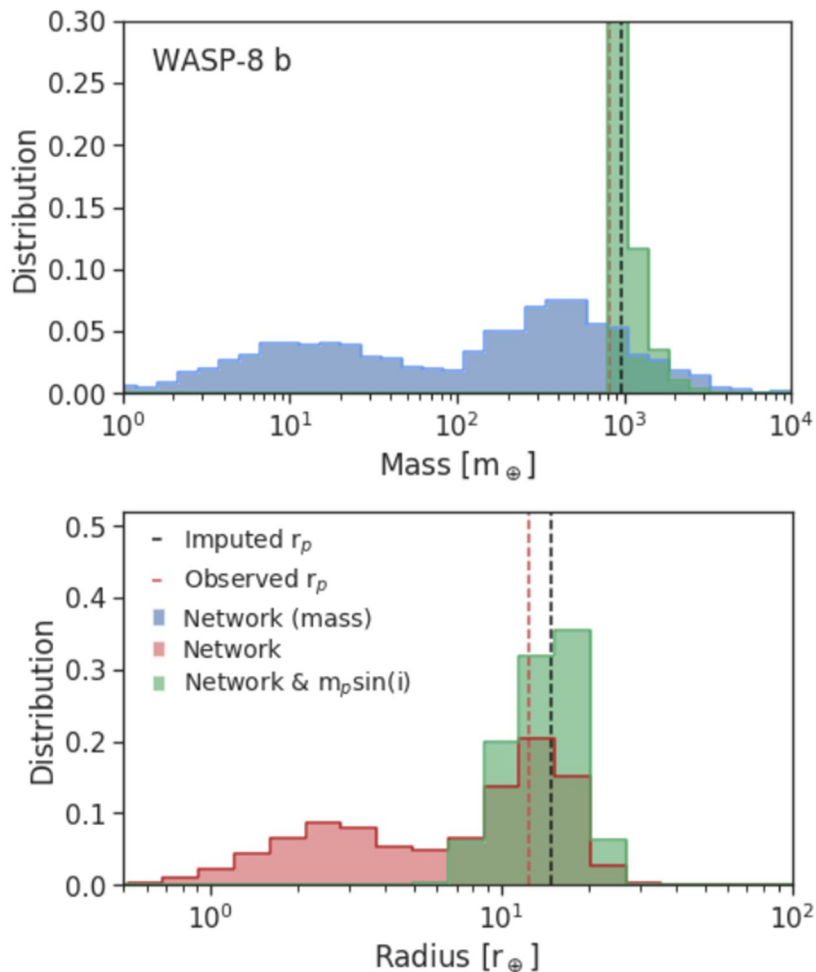


Figure 1: The neural network was tested for accuracy on planets with measured values. The top graph show the estimated distribution for the planet's mass, based on measurements for stellar mass, orbital period, number of planets in the system and effective temperature in blue. The green distribution is when a measurement for the minimum mass is included. Bottom graph is the same, but for the planet radius. The bimodal distribution suggests two possible planet classes commonly exist with similar orbits around similar stars.

Notably, both the mass and radius of this planet are bimodal. This suggests two distinct classes of planet co-exist in the same orbital space, rather than a single class of planet that may experience an inflated atmosphere. This points against the trend proposed in the original goal for this project that inflated atmospheres are common through evolution events such as planetesimal accretion.

WASP-8b is in truth a hot Jupiter, whose measured minimum mass could break the network degeneracy. Models for planet formation therefore need to allow these two distinct planetary groups to evolve to exist in the same orbital space around the star, despite the fact their formation and migration rates should differ.

(3) A challenge in the planetary science community is to explain the potential for planetary diversity within the exoplanet catalogue outside the field. Media and even journal papers frequently describe Earth-sized planets as "Earth-like" despite the presence of the near-equal sized Earth and Venus in our own Solar System.

To combat this, a simple climate model was developed to mimic the Earth's carbon cycle. The model was light enough to run on a website and users can experiment with changing the land

fraction, habitable zone position and volcanic rate and see how it impacts the average surface temperature. The goal is to demonstrate how even small changes in properties can make a big difference to a planet. The website is at earthlike.world. A neural network was used to render a possible image of the planet surface, based on its properties. This was published in the International Journal of Astrobiology.

(4) The image processing abilities of neural networks were utilised in an additional project to identify craters on the surface of Mars. Crater counting is commonly a tedious job performed by hand. The results are used to age the surfaces of celestial bodies. Neural networks can potentially accelerate the task by learning to recognize crater locations and sizes. This network has a success rate for identifying craters between 65% - 76% in common with human annotated data. This work was published in IEEE Journal of Selected Topics in Applied Earth Observations and Remote Sensing.

5. 主な発表論文等

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〔産業財産権〕

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6. 研究組織

	氏名 (ローマ字氏名) (研究者番号)	所属研究機関・部局・職 (機関番号)	備考
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